

5th International Workshop on
Dark Matter, Dark Energy and Matter-Antimatter Asymmetry

暗物質、暗能量及物質-反物質不對稱

December 28, 2018 - National Center for Theoretical Sciences, Hsinchu, Taiwan

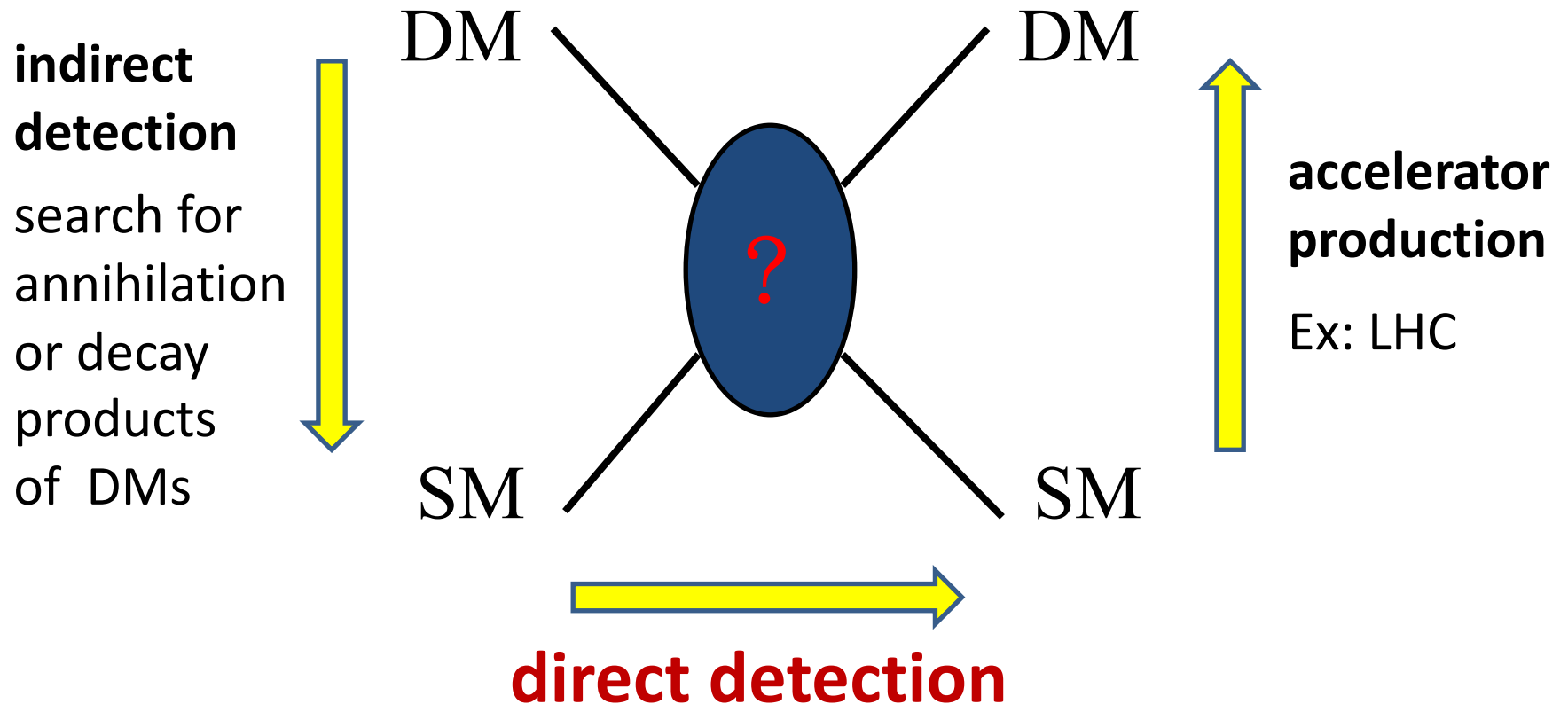
December 29-31, 2018 - Fo-Guang-Shan, Kaohsiung, Taiwan

Atomic corrections for direct detection of neutrinos and dark matters

Chih-Pan Wu

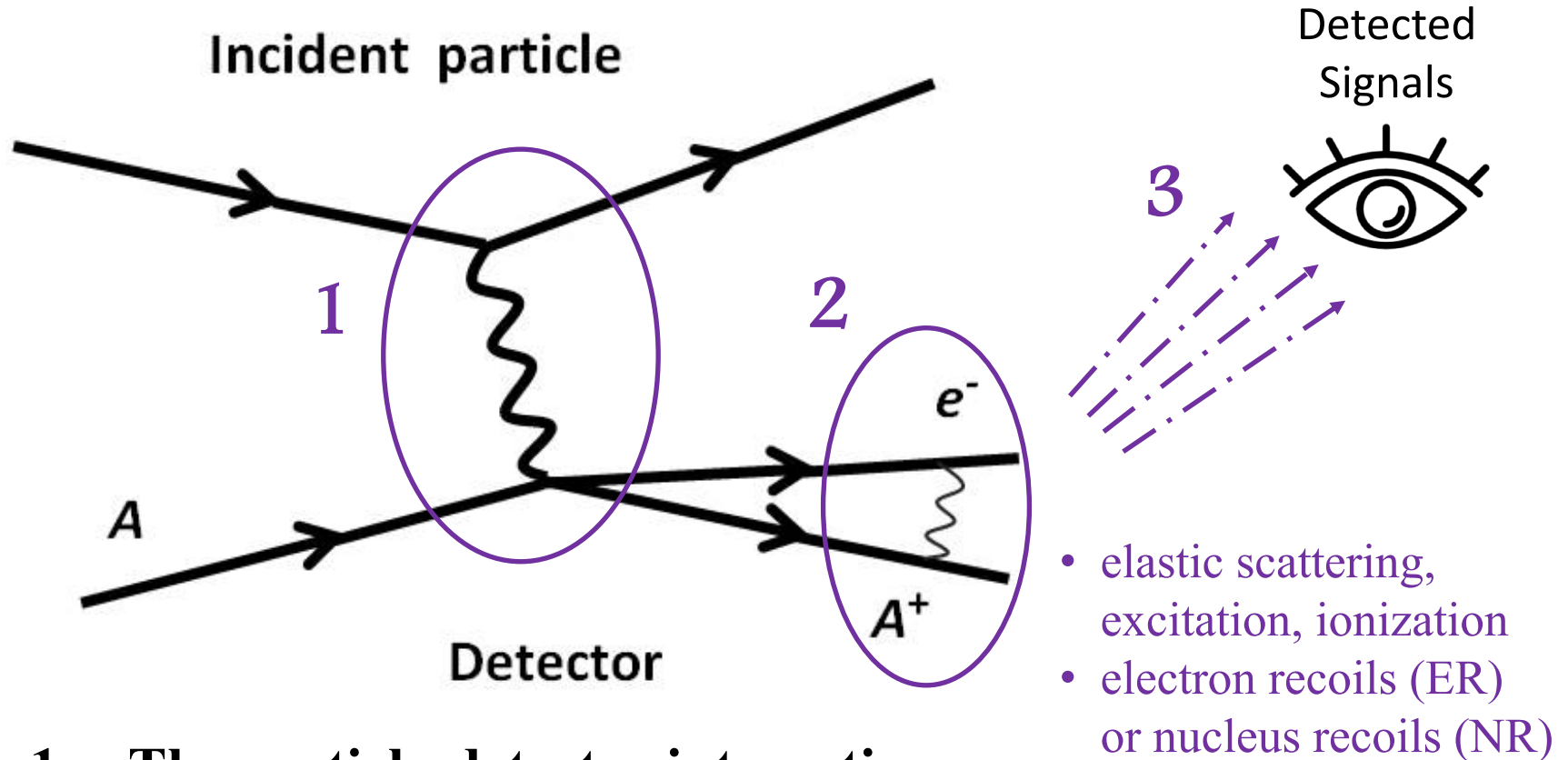
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(NCTS Experimental Collaboration Program 2)

Strategies: Search for Dark Matters



Using Semiconductor (Ex: high-purity Ge) or scintillator (Ex: liquid-Xe or crystal) to detect the existence of DM.

Scattering Diagrams and Detector Response



1. The particle-detector interaction
2. $d\sigma/dT$ for the primary scattering process
3. The following energy loss mechanism

DM Effective Interaction with Electron or Nucleons

Leading order (LO):

short range

$$L_{\text{int}}^{(\text{LO})} = \sum_{f=e,p,n} \left\{ \boxed{c_1^{(f)} (\chi^\dagger \chi) (f^\dagger f)} + \boxed{c_4^{(f)} (\chi^\dagger \vec{S}_\chi \chi) \cdot (f^\dagger \vec{S}_f f)} \right. \\ \left. + \boxed{d_1^{(f)} \frac{1}{q^2} (\chi^\dagger \chi) (f^\dagger f)} + \boxed{d_4^{(f)} \frac{1}{q^2} (\chi^\dagger \vec{S}_\chi \chi) \cdot (f^\dagger \vec{S}_f f)} \right\}$$

spin-indep. long range spin-dep.

Differential cross section for spin-independent contact interaction with electron ($c_1^{(e)}$) :

$$d\sigma|_{c_1^{(e)}} = \frac{2\pi}{v_\chi} \sum_F \sum_I |\langle F | \boxed{c_1^{(e)}} e^{i\frac{\mu}{m_e} \vec{q} \cdot \vec{r}} \boxed{|I\rangle}|^2 \quad \text{Initial \& final states of detector material}$$

$$\times \delta(T - E_{\text{c.m.}} - (E_F - E_I)) \frac{d^3 k_2}{(2\pi)^3}$$

Source : DM v.s. Neutrino

- Neutrino

- $m_\nu \rightarrow 0$, $E_\nu \sim k_\nu$ (few keV $\sim$ MeV)
- For given energy transfer T , 3-momentum transfer region:

$$T < Q < 2E_\nu - T \quad q^2 < 0$$

- Cold Dark Matter

- $m_\chi \gg m_e$, $E_\chi \sim 1/2 m_\chi v_\chi^2$, $k_\chi \sim m_\chi v_\chi$
- For given energy transfer T , 3-momentum transfer region:

$$m_\chi v_\chi - (m_\chi^2 v_\chi^2 - 2m_\chi T)^{1/2} \sim m_\chi v_\chi + (m_\chi^2 v_\chi^2 - 2m_\chi T)^{1/2}$$

$\gg$ outgoing electron momentum

because $2m_e T \ll 2m_\chi T < m_\chi^2 v_\chi^2$

$$q^2 \ll 0$$

Atomic Response Function

$$R_L \equiv \sum_{m_{jf}} \overline{\sum_{m_{ji}}} \int \frac{d^3 \vec{p}_r}{(2\pi)^3} |\langle f | \rho^{(A)}(\vec{q}) | i \rangle|^2 \delta\left(T - B - \frac{q^2}{2M} - \frac{p_r^2}{2\mu_{\text{red}}}\right)$$

Continuous WF:

$$\propto \text{Exp}[i p_r r]$$

$$p_r = (2m_e T)^{1/2}$$

$$\rho^{(e)}(\vec{q}) = e^{i \frac{\mu}{m_e} \vec{q} \cdot \vec{r}}$$

Highly oscillated

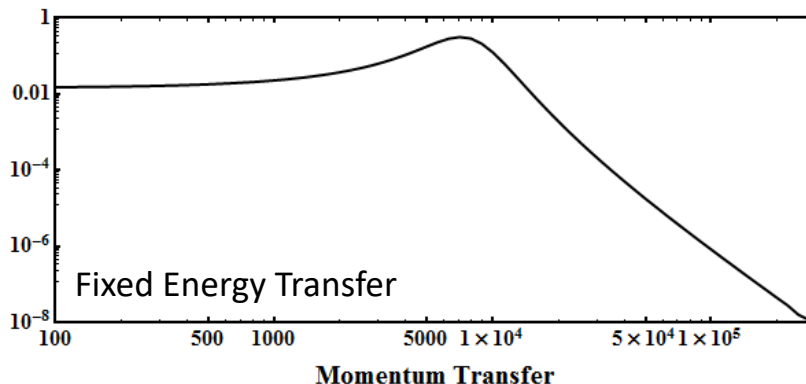
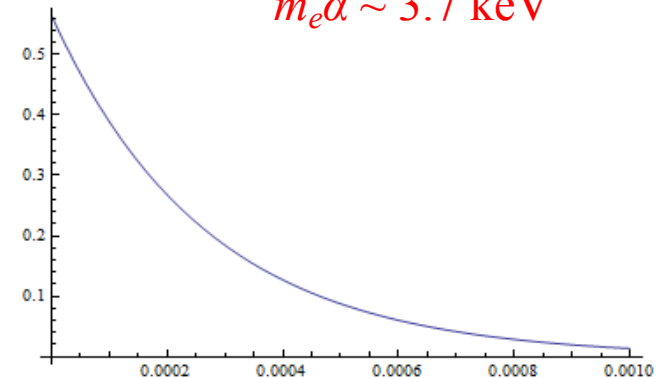
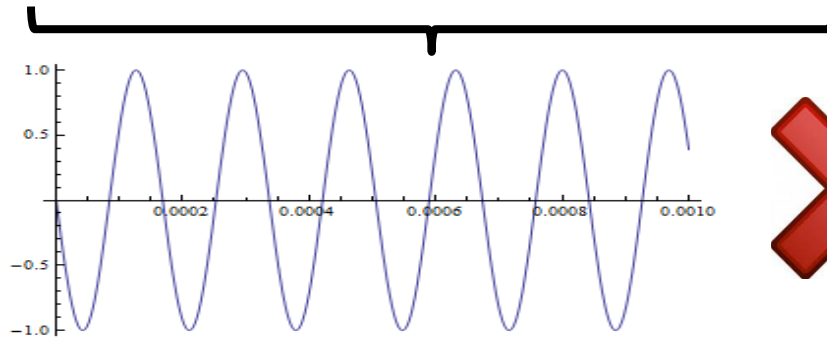
$$\rho^{(p)}(\vec{q}) = e^{-i \frac{\mu}{m_p} \vec{q} \cdot \vec{r}}$$

1/2000 smaller

bound WF:

$$\propto \text{Exp}[-Z m_e \alpha r]$$

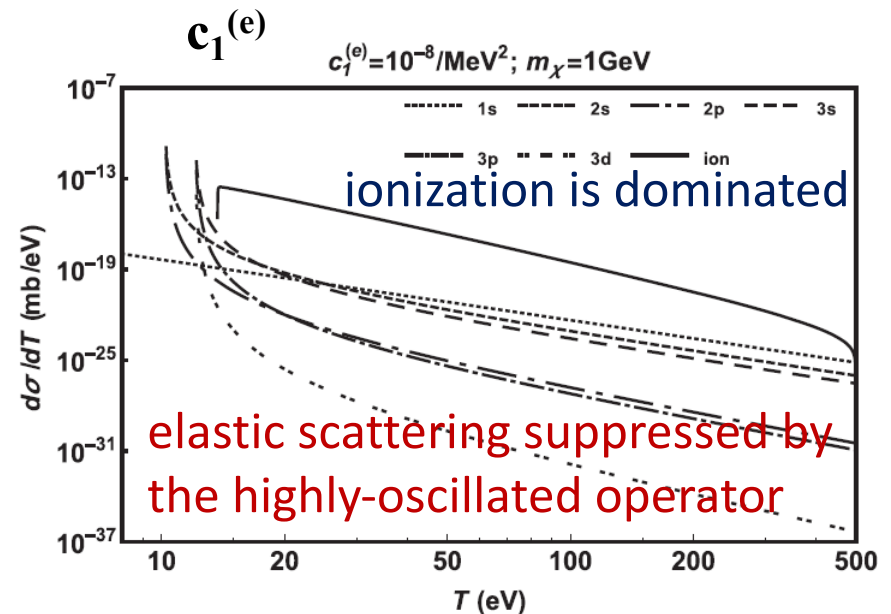
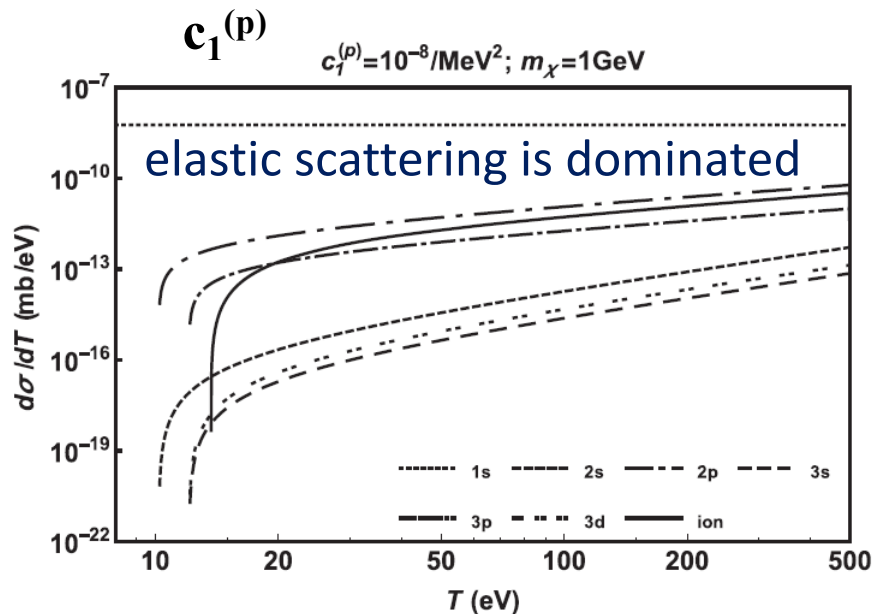
$$m_e \alpha \sim 3.7 \text{ keV}$$



Maximum contribution
when $p_r \sim \mu/m q \gg Z m_e \alpha$

NR v.s. ER : Hydrogen Toy Model

Spin-independent contact interaction with proton and electron



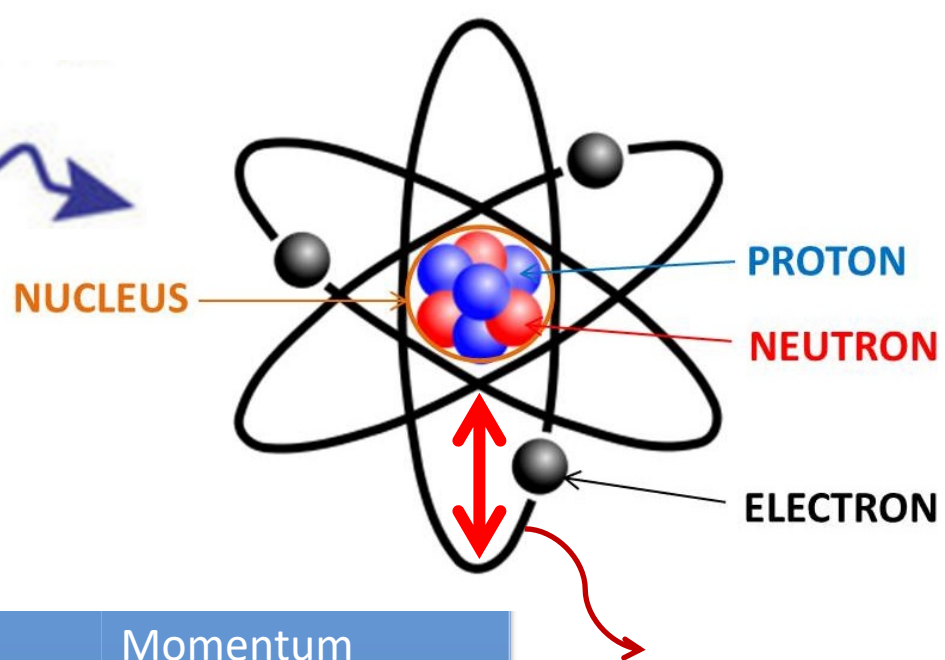
But in free electron case, this suppression doesn't exist, will be wrong for interactions with electron.

➡ Free electron assumption will be several orders of magnitude over estimation for ER

Why we study Atomic Response ?

The space uncertainty is inversely proportional to its incident momentum:

$$\lambda \sim 1/p$$



Atomic size is related to its orbital momentum:

$$Z m_e \alpha \sim Z * 3.7 \text{ keV}$$

Z: effective charge

LDM with velocity $\sim 10^{-3}$

Mass	Energy	Momentum
1 GeV	$m_\chi + 500 \text{ eV}$	1 MeV
100 MeV	$m_\chi + 50 \text{ eV}$	100 keV
Neutrino Sources		
Reactor ν	$\sim \text{few MeV}$	Same as energy
Solar ν (pp)	$\sim \text{few hundred keV}$	Same as energy

When atomic structures should be considered (free target approx. fail)?

- Incident momentum ~ 100 keV and below
 - The wavelengths of incident particles are about the same order with the size of the atom.
 - For Innermost orbital, the related momentum $\sim Z m_e \alpha \sim Z * 3$ keV (Z = effective nuclear charge)
- Energy transfer ~ 10 keV and below
 - barely overcome the atomic thresholds
 - For Innermost orbital, binding energy ~ 11 keV (Ge) and 34 keV (Xe)
- Strong phase-space suppression for DM-e scattering

Ab initio Theory for Atomic States

MCDF: multiconfiguration Dirac-Fock method

Dirac-Fock method: $\psi(t)$ is a Slater determinant of one-electron orbitals $u_a(\vec{r}, t)$ and invoke variational principle $\delta \langle \bar{\psi}(t) | i \frac{\partial}{\partial t} - H - V_I(t) | \psi(t) \rangle = 0$ to obtain eigenequations for $u_a(\vec{r}, t)$.

multiconfiguration: Approximate the many-body wave function $\Psi(t)$ by a superposition of configuration functions $\psi_\alpha(t)$

$$\Psi(t) = \sum_{\alpha} C_{\alpha}(t) \psi_{\alpha}(t)$$

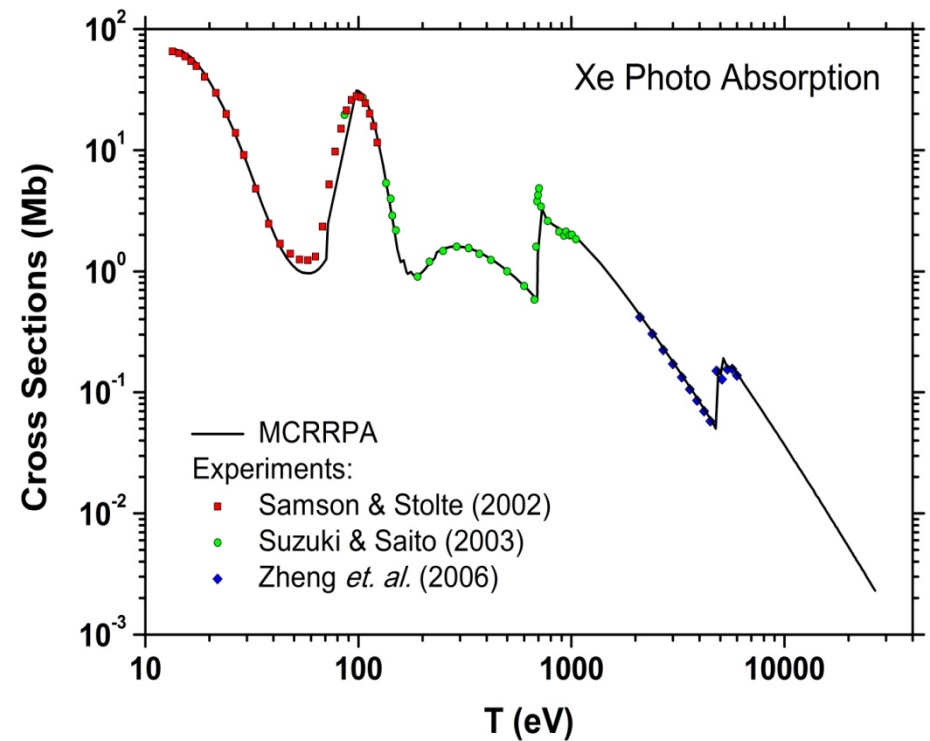
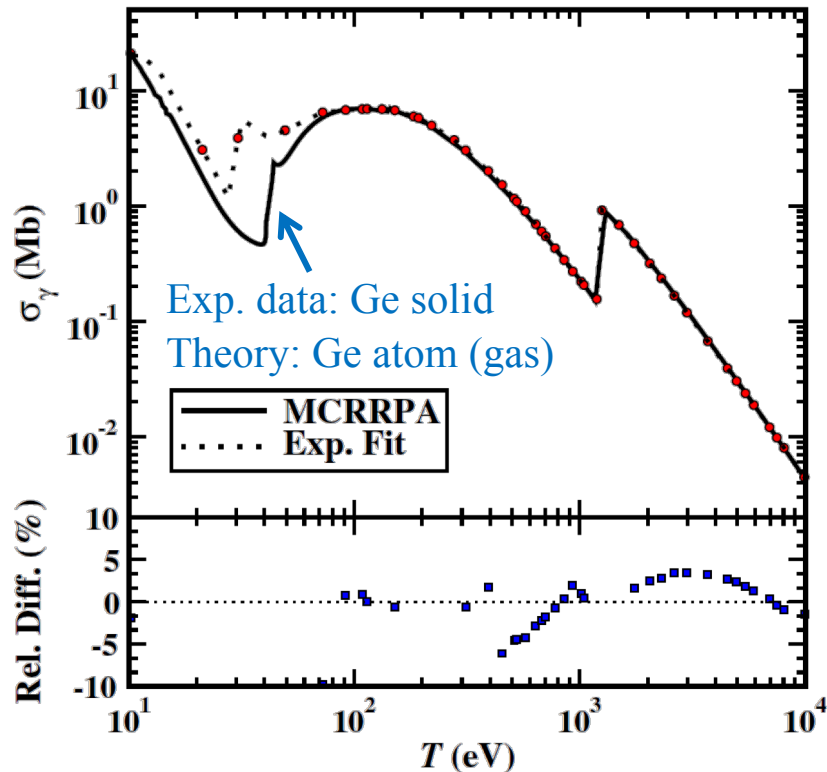
MCRRPA: multiconfiguration relativistic random phase approximation

RPA: Expand $u_a(\vec{r}, t)$ into time-indep. orbitals in power of external potential

$$u_a(\vec{r}, t) = e^{i\varepsilon_a t} \left[u_a(\vec{r}) + w_{a+}(\vec{r}) e^{-i\omega t} + w_{a-}(\vec{r}) e^{i\omega t} + \dots \right]$$

$$C_a(t) = C_a + [C_a]_+ e^{-i\omega t} + [C_a]_- e^{i\omega t} + \dots$$

Benchmark : Ge & Xe Photoionization



Above 100 eV error under 5%.

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J. Samson and W. Stolte, J. Electron Spectrosc. Relat. Phenom. **123**, 265 (2002).
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Recent Outputs

- Constraints on millicharged particles with low threshold germanium detectors at Kuo-Sheng Reactor Neutrino Laboratory
[@Lakhwinder Singh, 11:05 Dec. 29](#)
- Constraints on spin-independent dark matter scattering off electrons with germanium and xenon detectors [@Mukesh Kumar Pandey, 12:05 Dec. 30](#)
- Discovery Power of Multi-ton Xenon Detectors in Light Dark Matter Searches and Exotic Neutrino Properties [@Chung-Chun Hsieh, 14:45 Dec. 31](#)

Thanks for your attention !!

Related Talks:

[Lakhwinder Singh, 11:05 Dec. 29](#)

[Mukesh K. Pandey, 12:05 Dec. 30](#)

[Chung-Chun Hsieh, 14:45 Dec. 31](#)

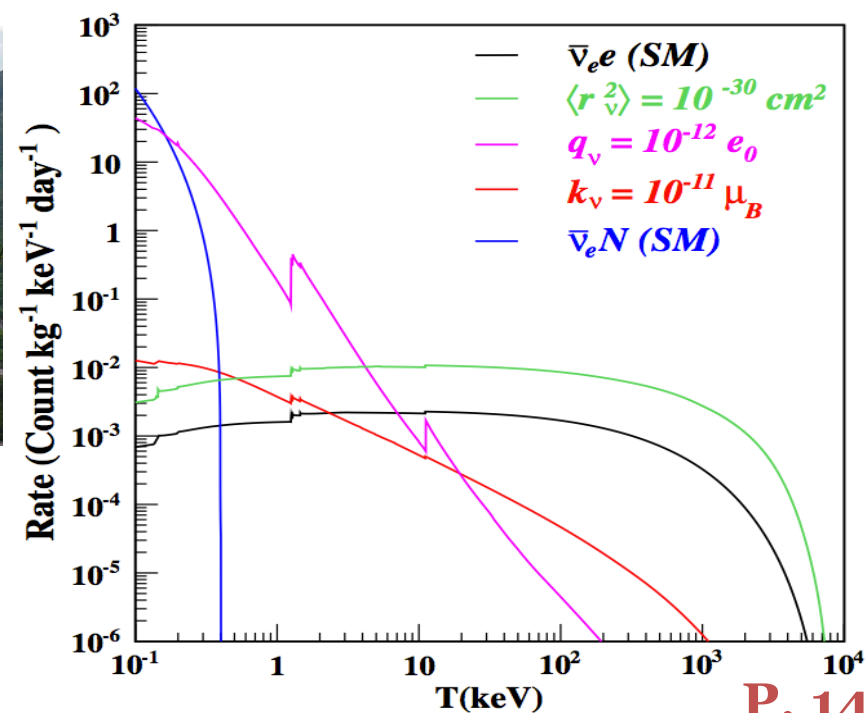
Applications I: Neutrino EM Properties

Data set	Reactor- $\bar{\nu}_e$	Data strength Reactor on/off (kg-days)	Analysis Threshold (keV)	Bounds at 90% C.L.		
	Flux ($\times 10^{13} \text{ cm}^{-2} \text{ s}^{-1}$)			$\kappa_{\bar{\nu}_e}^{(\text{eff})}$ ($\times 10^{-11} \mu_B$)	$q_{\bar{\nu}_e}$ ($\times 10^{-12}$)	$\langle r_{\bar{\nu}_e}^2 \rangle^{(\text{eff})}$ ($\times 10^{-30} \text{ cm}^2$)
TEXONO 187 kg CsI [9]	0.64	29882.0/7369.0	3000	< 22.0	< 170	< 0.033
TEXONO 1 kg Ge [5,6]	0.64	570.7/127.8	12	< 7.4	< 8.8	< 1.40
GEMMA 1.5 kg Ge [7,8]	2.7	1133.4/280.4	2.8	< 2.9	< 1.1	< 0.80
TEXONO point-contact Ge [4,17]	0.64	124.2/70.3	0.3	< 26.0	< 2.1	< 3.20
Projected point-contact Ge	2.7	800/200	0.1	< 1.7	< 0.06	< 0.74
Sensitivity at 1% of SM	...	...	...	~ 0.023	~ 0.0004	~ 0.0014



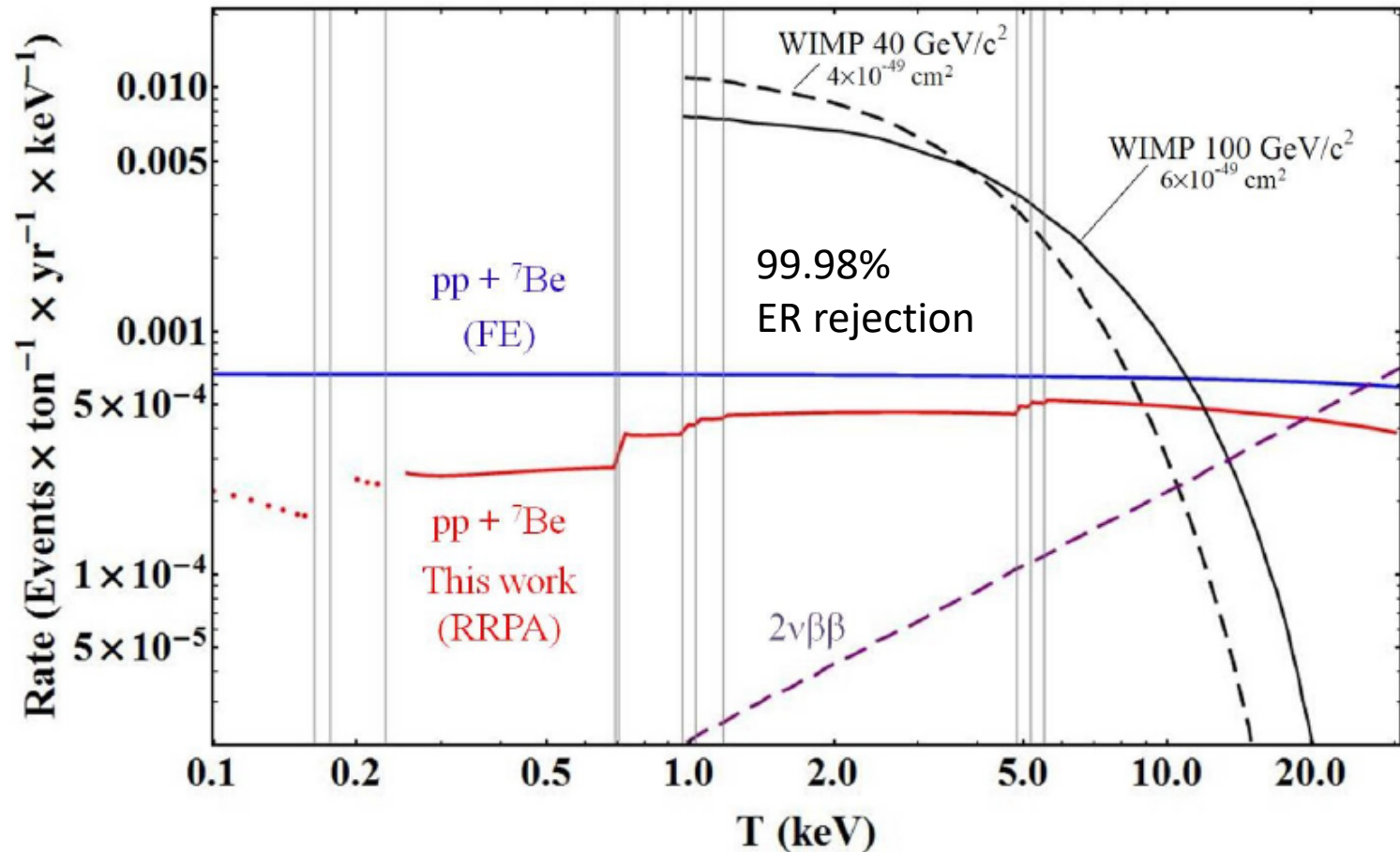
Reference:

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Applications II:

Solar ν Background in LXe Detectors

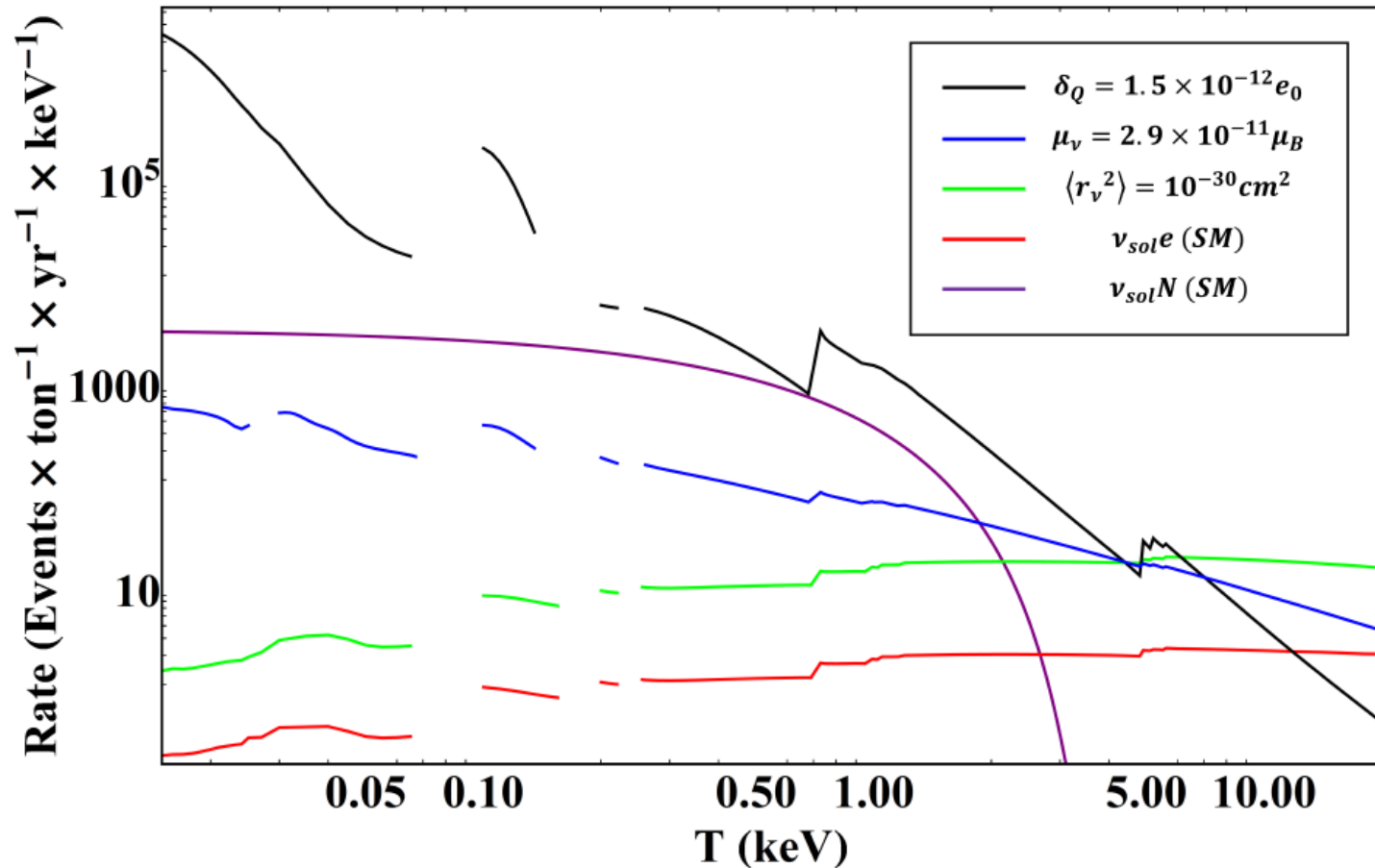


J. Aalbers *et. al.* (DARWIN collaboration), arXiv:1606.07001 (2016).

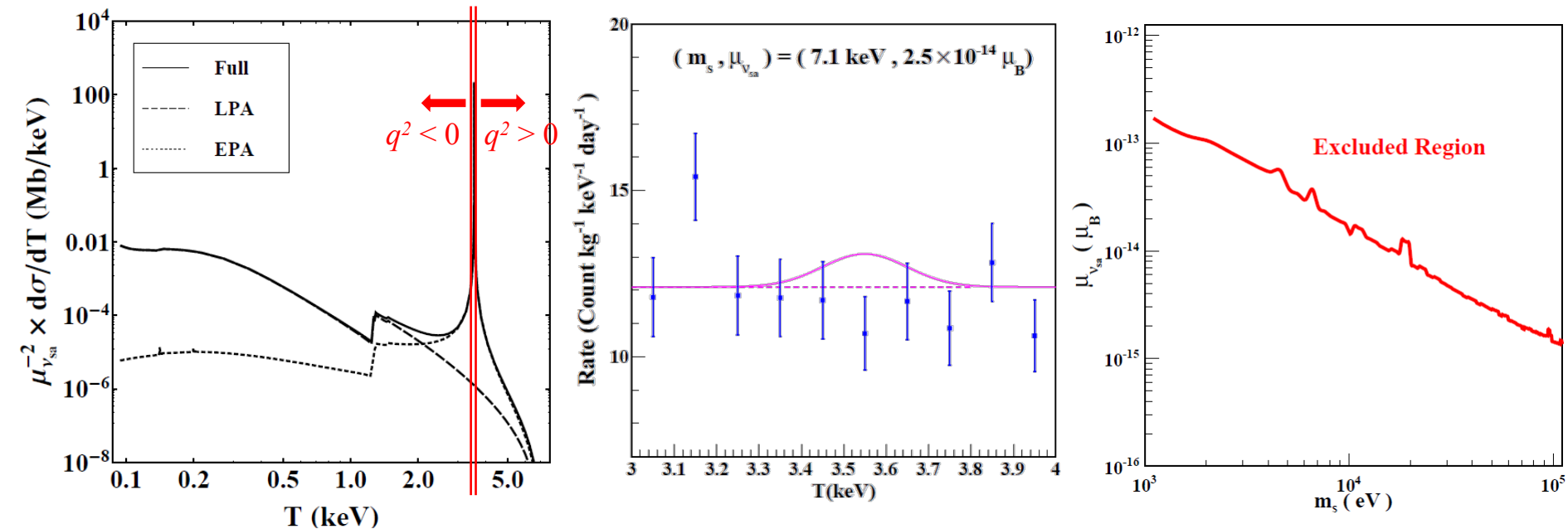
J.-W. Chen *et. al.*, arXiv:1610.04177 (2016).

Applications II-2:

Solar ν As Signals in LXe Detectors



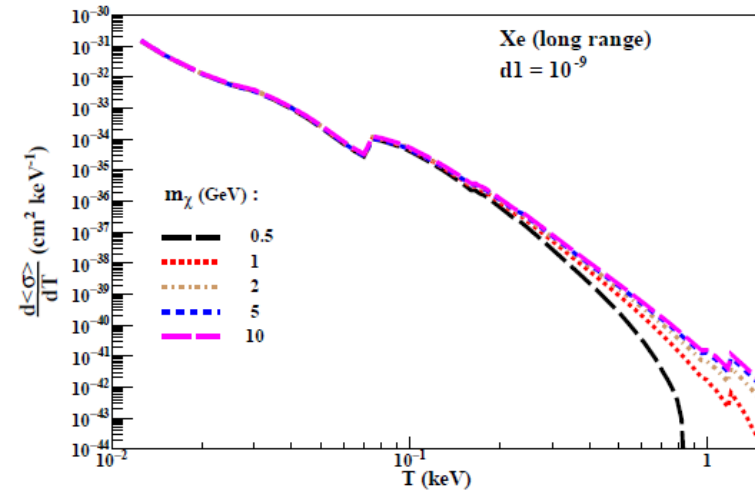
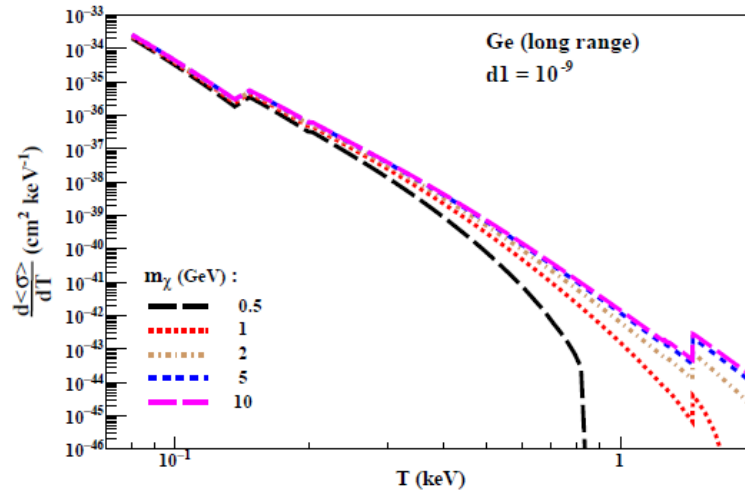
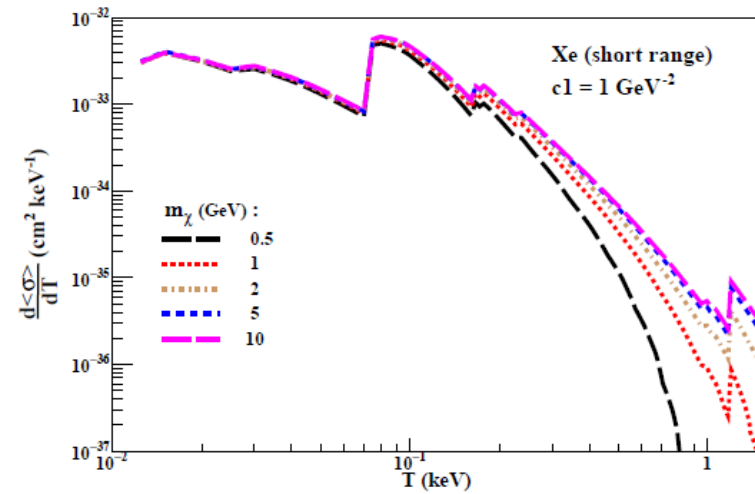
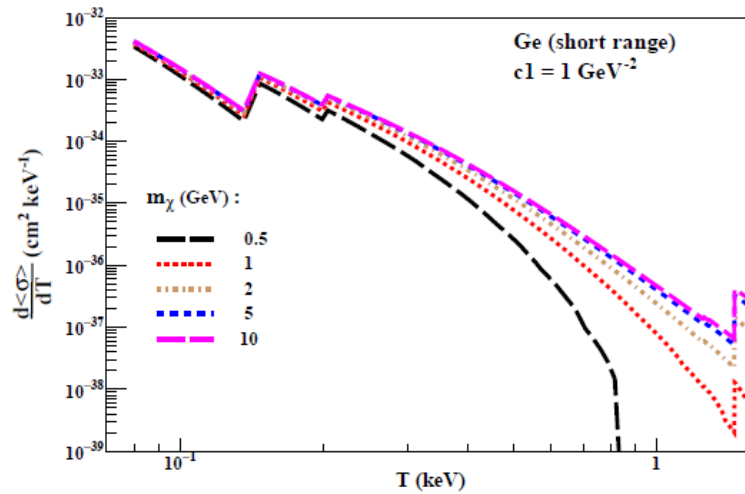
Applications III: Sterile Neutrino Direct Constraint



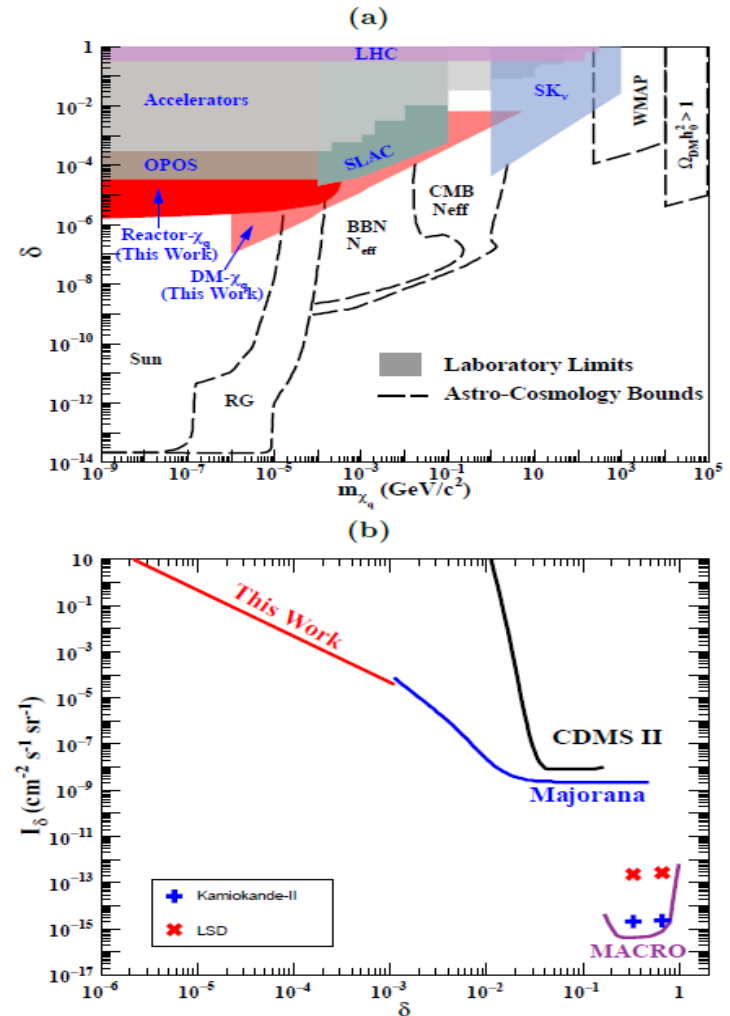
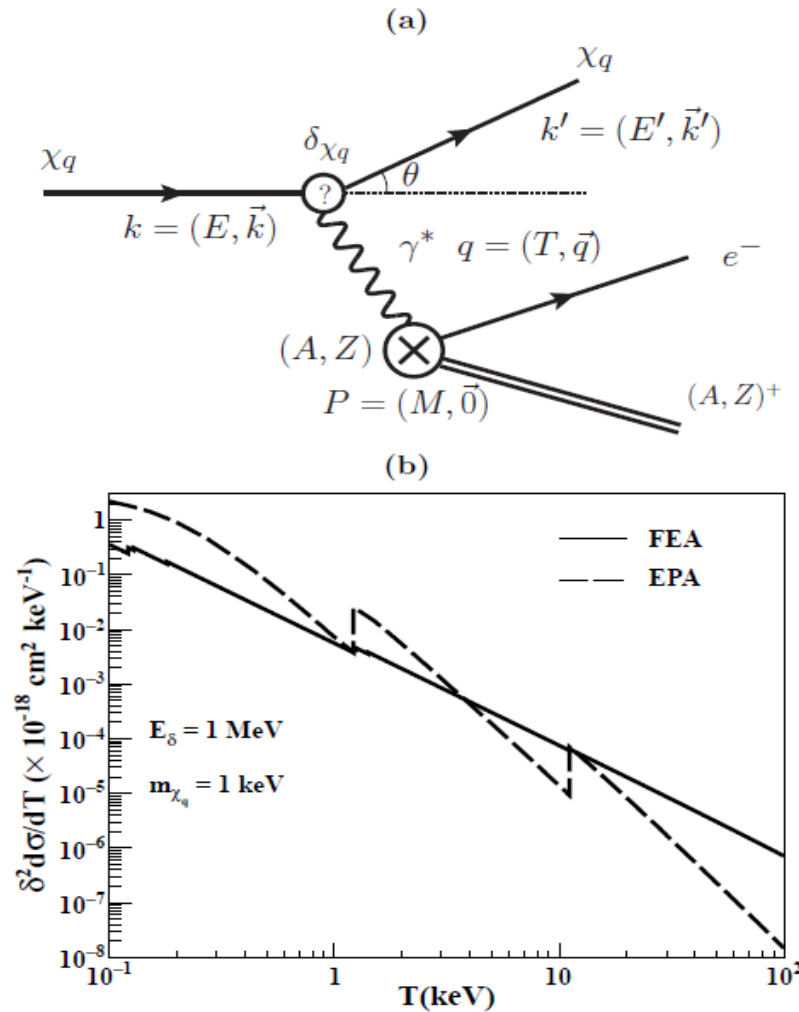
- Non-relativistic massive sterile neutrinos decay into SM neutrino.
- At $m_s = 7.1 \text{ keV}$, the upper limit of $\mu_{\nu_{sa}} < 2.5 \times 10^{-14} \mu_B$ at 90% C.L.
- The recent X-ray observations of a 7.1 keV sterile neutrino with decay lifetime $1.74 \times 10^{-28} \text{ s}^{-1}$ can be converted to $\mu_{\nu_{sa}} = 2.9 \times 10^{-21} \mu_B$, much tighter because its much larger collecting volume.

Applications IV:

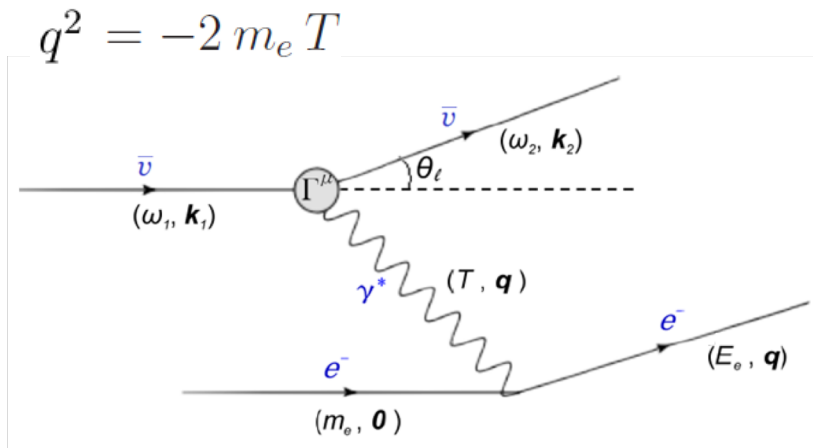
Spin-Indep. DM-e Scattering in Ge & Xe



Applications V: Constraints on millicharged DM particles

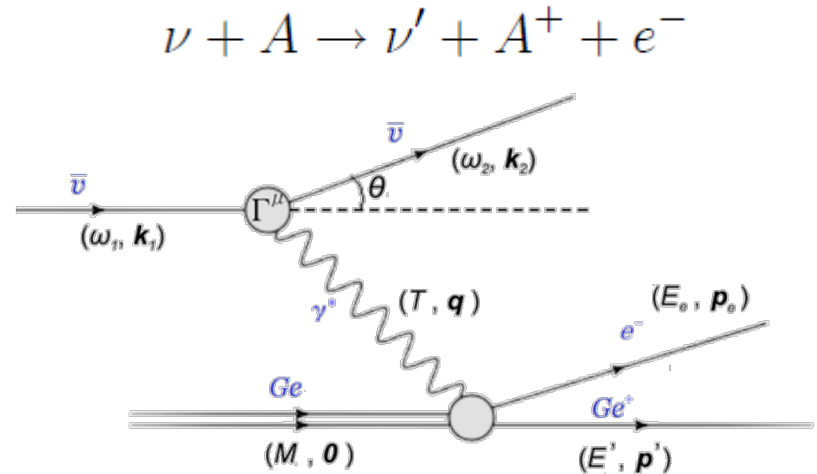


Target: Free e/n v.s. Atom



Phase space is fixed in 2-body scattering
 → 4-momentum transfer is fixed
 → scattering angle is fixed
 → Maximum energy transfer is limited

by a factor
$$r = \frac{4 m_{inc} m_{tar}}{(m_{inc} + m_{tar})^2}$$



Energy and momentum transfer can be shared by nucleus and electrons
 → Inelastic scattering
 (energy loss in atomic energy level)
 → Phase space suppression

Toy Model : Analytic Hydrogen WFs

$$\langle 100 | \vec{r} \rangle = \frac{1}{\sqrt{\pi}} Z^{\frac{3}{2}} e^{-Z\bar{r}}, \quad \text{exp.-decay with the rate } \propto \text{orbital momentum} \sim 3.7 \text{ keV}$$

$$\langle nlm_l | \vec{r} \rangle = \frac{1}{(2l+1)!} \sqrt{\frac{(n+l)!}{2n(n-l-1)!}} \left(\frac{2Z}{n}\right)^{\frac{3}{2}} e^{-\frac{Z\bar{r}}{n}} \left(\frac{2Z\bar{r}}{n}\right)^l {}_1F_1\left(-(n-l-1), 2l+2, \frac{2Z\bar{r}}{n}\right) Y_l^{m_l*}(\theta, \phi),$$

$$\langle \vec{p}_r | \vec{r} \rangle = e^{\frac{\pi Z}{2\bar{p}_r}} \Gamma\left(1 - \frac{iZ}{\bar{p}_r}\right) e^{-i\vec{p}_r \cdot \vec{r}} {}_1F_1\left(\frac{iZ}{\bar{p}_r}, 1, i(p_r r + \vec{p}_r \cdot \vec{r})\right)$$

Oscillated like sin/cos function with frequency $\propto$ electron momentum $\sim (2m_e T)^{1/2}$

- The initial state of the hydrogen atom at the ground state, the spatial part $|I\rangle_{\text{spat}} = |1s\rangle$
- 1. elastic scattering:** $\langle F |_{\text{spat}} = \langle 1s |$
- 2. discrete excitation (ex):** $\langle F |_{\text{spat}} = \langle nlm_l |$
- 3. ionization (ion):** $\langle F |_{\text{spat}} = \langle \vec{p}_r |$

Reference:

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Thanks for your attention!